

Expanding impulsive gravitational waves

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Outline

1 Motivation

General relativity

Gravitational waves: real physics!

Explicit (unrealistic) toy-models

2 Expanding impulsive waves

Spacetime geometry

Geodesics

3 Sources of expanding impulses

Complex mapping $h(Z)$

One string

Two strings

4 Conclusions



General relativity: Einstein's theory of gravity

Gravity – **universal interaction** – inherent property of the ‘**arena**’ (spacetime)

- natural application of differential geometry: spacetime \leftrightarrow Lorentzian manifold
- WANTED: metric tensor g_{ab} (dead or alive)

Einstein's field equations

$$R_{ab} - \frac{1}{2}R g_{ab} + \Lambda g_{ab} = \frac{8\pi G}{c^4} T_{ab}$$

‘ *geometry = energy and momentum* ’



(credit: ligo.org)

Non-linearity: *matter* \rightarrow *space curvature* \rightarrow *matter motion* \rightarrow *curvature changed* \rightarrow **etc.**

Corollary: ripples in the spacetime curvature propagating with the speed of light
 \leftrightarrow **gravitational waves**

Solutions: exact spacetimes \times perturbative models \times numerical simulations

Free fall: geodesic motion $\leftrightarrow \ddot{x}^a + \Gamma_{bc}^a \dot{x}^b \dot{x}^c = 0$



Gravitational waves: Sci-fi? No!

February 11, 2016, Washington, D.C.

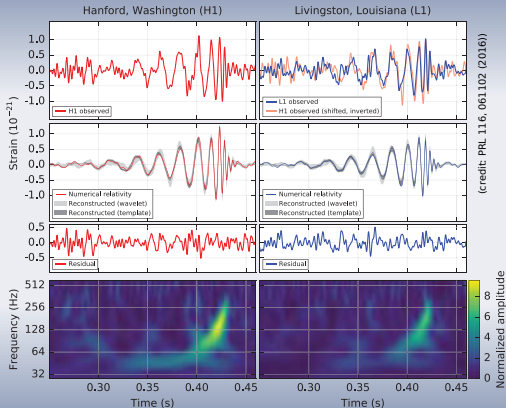
Ladies and gentlemen,
We have detected gravitational waves!
We did it!



David Reitze (LIGO Executive Director)



Rainer Weiss and Kip Thorne



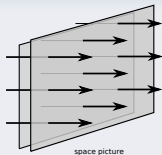
event: **GW150914**
detected: **LIGO**, September 14, 2015, 09:50:45 UTC
lag: **6,9 ms H1** behind **L1**



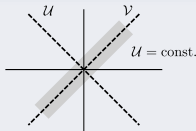
Exact toy-models: Sandwich waves

Flat Minkowski space + 'something': $ds^2 = -dt^2 + dx^2 + dy^2 + dz^2 + \text{'something'}$

- *Non-expanding:* $\mathcal{U} = \frac{1}{\sqrt{2}}(t - z)$ and $\mathcal{V} = \frac{1}{\sqrt{2}}(t + z) \rightarrow$
 $ds^2 = dx^2 + dy^2 - 2d\mathcal{U}d\mathcal{V} + \text{'something'}$

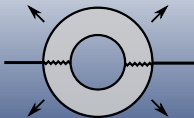


space picture

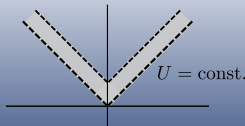


spacetime picture

- *Expanding:* $\mathcal{U} = Z\bar{Z}V - U$, $\mathcal{V} = V$, $\eta = ZV$ where $\eta = \frac{1}{\sqrt{2}}(x + iy) \rightarrow$
 $ds^2 = 2V^2 dZd\bar{Z} + 2dUdV + \text{'something'}$



space picture



spacetime picture

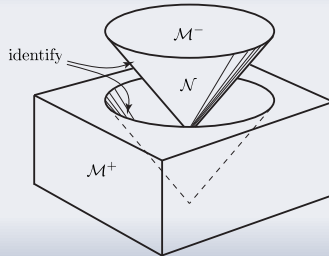


Cut and paste construction of expanding impulses

Minkowski space in the null cone foliation:

$$ds^2 = 2 \frac{V^2}{|1 + \epsilon Z \bar{Z}|^2} dZ d\bar{Z} + 2 dU dV - 2\epsilon dU^2$$

Cut along the null cone:



Identification via the Penrose junction condition:

$$[U = 0_-, V, Z, \bar{Z}]_{\mathcal{M}^-} \equiv \left[U = 0_+, \frac{1 + \epsilon h \bar{h}}{1 + \epsilon Z \bar{Z}} \frac{V}{|h'|}, h(Z), \bar{h}(\bar{Z}) \right]_{\mathcal{M}^+}$$



Continuous line element

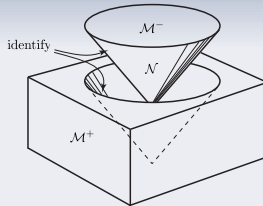
Flat Minkowski space:

$$ds^2 = -2 d\mathcal{U} d\mathcal{V} + 2 d\eta d\bar{\eta}$$

Explicit cut and paste procedure:

- for $U < 0$

$$\begin{aligned} \mathcal{U} &= \frac{Z\bar{Z}}{p} V - U \\ \mathcal{V} &= \frac{V}{p} - \epsilon U \\ \eta &= \frac{Z}{p} V \end{aligned}$$



- for $U > 0$

$$\begin{aligned} \mathcal{U} &= BV - EU \\ \mathcal{V} &= AV - DU \\ \eta &= CV - FU \end{aligned}$$

Resulting metric: \mathcal{M}^- and \mathcal{M}^+ re-attached along **null cone** $U = 0$ with a ‘warp’

$$ds^2 = 2 \left| \frac{V}{p} dZ + U_+(U) p \bar{H} d\bar{Z} \right|^2 + 2 dU dV - 2\epsilon dU^2$$

where

$$U_+ \equiv U_+(U) = \begin{cases} 0 & \text{if } U \leq 0 \\ U & \text{if } U \geq 0 \end{cases} \quad \text{and} \quad H(Z) = \frac{1}{2} \left[\frac{h'''}{h'} - \frac{3}{2} \left(\frac{h''}{h'} \right)^2 \right]$$



Basic properties of expanding impulses

Continuous metric:

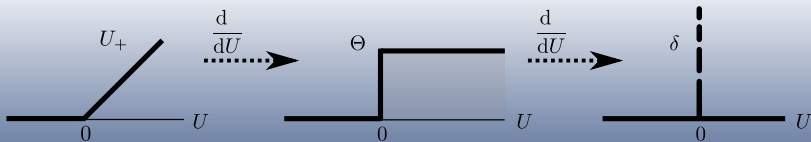
$$ds^2 = 2 \left| \frac{V}{p} dZ + U_+(U) p \bar{H} d\bar{Z} \right|^2 + 2 dU dV - 2\epsilon dU^2$$

Penrose junction conditions: satisfied on $U = 0$

Curvature: discontinuity in the derivatives of the metric \rightarrow impulsive components

- Ricci tensor: $\Phi_{22} = \frac{v^4 H \bar{H}}{v^2} U \delta(U)$
- Weyl tensor: $\Psi_4 = \frac{v^2 H}{V} \delta(U)$

where $p = 1 + \epsilon Z \bar{Z}$ and $\epsilon = -1, 0, +1$



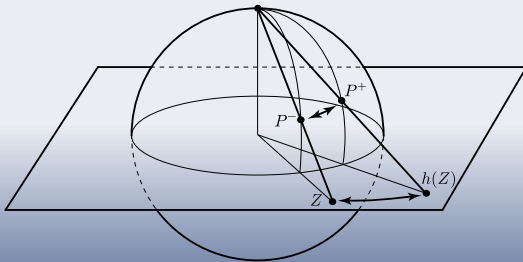


Geometric nature of junction condition

Penrose junction condition: geometric nature

$$[U = 0_-, V, Z, \bar{Z}]_{\mathcal{M}^-} \equiv [U = 0_+, \frac{1+\epsilon h\bar{h}}{1+\epsilon Z\bar{Z}} \frac{V}{|h'|}, h(Z), \bar{h}(\bar{Z})]_{\mathcal{M}^+}$$

- the impulse on $U = 0 \leftrightarrow$ sphere expanding with the speed of light: $x^2 + y^2 + z^2 = t^2$
- $h(Z)$ is an arbitrary *holomorphic function*
- identification of points in complex plane $Z \rightarrow h(Z) \leftrightarrow$ *stereographic projection*





Geodesics: global uniqueness and C^1 -regularity

J. Podolský, C. Sämann, R. Steinbauer, R. Švarc: *The global uniqueness and C^1 -regularity of geodesics in expanding impulsive gravitational waves*, Class. Quantum Grav. **33** (2016) 195010

Theorem: *Existence and Uniqueness.*

For the entire class of expanding impulsive waves on any background of constant curvature described by the continuous form of the metric with smooth H we have: Given any point \mathcal{P} and any direction $v \in T_{\mathcal{P}}M$ there exists a *unique* C^1 -solution γ in the sense of Filippov to the geodesic equations with this initial data.

Moreover, if such a geodesic meets the impulsive wave located at $\mathcal{N} = \{U = 0\}$ at all, it is either one of its null generators or it hits it in *isolated points*.

Corollary: *Preservation of causal character.*

The geodesics γ of the theorem above satisfy $g(\dot{\gamma}, \dot{\gamma}) = \text{const.}$, and, in particular, the causal character of γ can be defined globally.

Corollary: *Crossing the expanding impulse.*

The geodesics of the theorem above that start off the wave surface $\mathcal{N} = \{U = 0\}$ and hit it at all, do so in isolated points either

- *transversally* and pass from the ‘outside’ to the ‘inside’, or vice versa, or
- *tangentially*, in which case they are spacelike and come from the ‘outside’ and revert to the ‘outside’ again.

Remark: *Uniqueness for non-smooth H*

If the function H has singularities then, given arbitrary initial data, the geodesic equation possesses locally defined unique C^1 -solutions in any region where H is sufficiently smooth.



Background coordinates: matching conditions on $U = 0$

Geodesics are straight lines outside the impulse: *matching across the impulse needed*

- begin with unique globally defined \mathcal{C}^1 -geodesics in the continuous coordinates
- evaluate the positions and velocities at the instant of interaction $U_i = U(\tau_i) = 0$
- apply transformations **in front of** and **behind** the impulse

Refraction formulae: *background geodesics crossing the spherical impulse*

- positions:

$$\begin{aligned}x_i^- &= |h'| \frac{Z_i + \bar{Z}_i}{h + \bar{h}} x_i^+ & y_i^- &= |h'| \frac{Z_i - \bar{Z}_i}{h - \bar{h}} y_i^+ \\z_i^- &= |h'| \frac{Z_i \bar{Z}_i - 1}{|h|^2 - 1} z_i^+ & t_i^- &= |h'| \frac{Z_i \bar{Z}_i + 1}{|h|^2 + 1} t_i^+\end{aligned}$$

- velocities:

$$\begin{aligned}\dot{x}_i^- &= a_x \dot{x}_i^+ + b_x \dot{y}_i^+ + c_x \dot{z}_i^+ + d_x \dot{t}_i^+ \\ \dot{y}_i^- &= a_y \dot{x}_i^+ + b_y \dot{y}_i^+ + c_y \dot{z}_i^+ + d_y \dot{t}_i^+ \\ \dot{z}_i^- &= a_z \dot{x}_i^+ + b_z \dot{y}_i^+ + c_z \dot{z}_i^+ + d_z \dot{t}_i^+ \\ \dot{t}_i^- &= a_t \dot{x}_i^+ + b_t \dot{y}_i^+ + c_t \dot{z}_i^+ + d_t \dot{t}_i^+\end{aligned}$$

where the coefficients a, b, c, d are complicated functions of $Z_i, h(Z_i)$ and its derivatives



Construction of a specific impulse

Junction condition: $h(Z)$ vs. **Curvature:** $H(Z)$

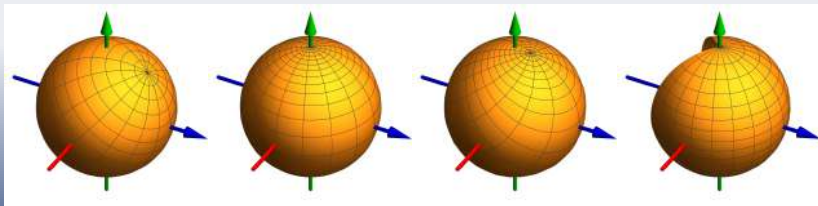
Möbius transformation: leaves $H(Z)$ unchanged

$$h(Z) : Z \mapsto \frac{aZ + b}{cZ + d}$$

Beyond linear fractions \mapsto *non-trivial Schwarzian derivative* $H(Z) \mapsto$ *topological defects*

Decomposition of $h(Z)$: elementary operations

- spatial rotations $\mathcal{R}_{\{\phi, \theta, \psi\}}Z$ parameterized by the Euler angles $\{\phi, \theta, \psi\}$
- boost in the z direction $\mathcal{B}_{\{w\}}Z$ parameterized by w
- the simplest string-like structure $\mathcal{S}_{\{\delta\}}Z = Z^{1-\delta}$: cuts out the wedge $2\pi\delta$ around z -axis



$$\mathcal{R}_{\{0, -\frac{\pi}{3}, 0\}}(Z)$$

$$\mathcal{B}_{\{1/3\}}Z$$

$$\mathcal{B}_{\{1/3\}}\mathcal{R}_{\{0, -\frac{\pi}{3}, 0\}}(Z)$$

$$\mathcal{S}_{\{1/4\}}Z$$

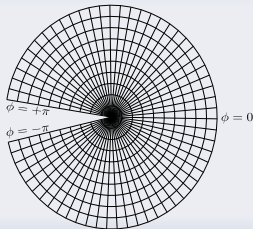


Specific function $h(Z)$: one string

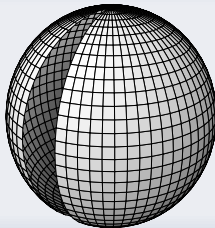
Assume: *complex mapping*

$$h(Z) = Z^{1-\delta}$$

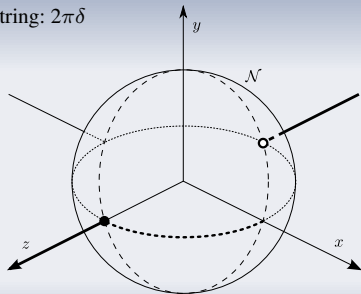
where $\delta \in [0, 1)$ characterizes *deficit angle* induced by a string: $2\pi\delta$



complex plane



Riemann sphere



space picture

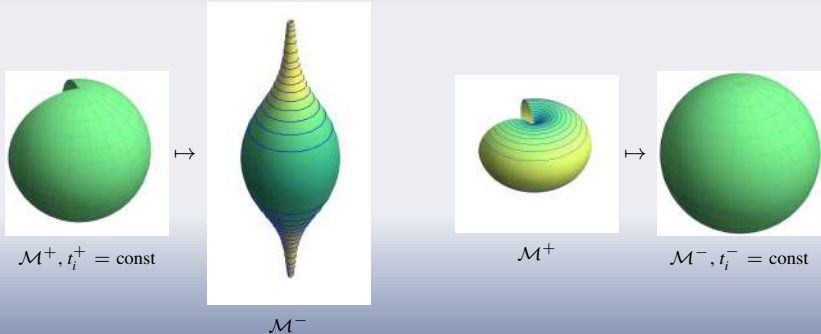
Metric function $H(Z)$: *Schwarzian derivative* of $h(Z)$ i.e. $H(Z) = \frac{\frac{1}{2}\delta(1-\frac{1}{2}\delta)}{Z^2}$



Specific function $h(Z)$: one string – initial data

Initial data for geodesics: static observers in front of the impulse

- interacting simultaneously at $t_i^+ = \text{const}$ \mapsto emerging at different t_i^-
- interacting at specific $t_i^+ =$ \mapsto emerging simultaneously at $t_i^- = \text{const}$



Colour changes show the *time shift* and the *lines of constant t_i^\pm* are thick blue lines.



Specific function $h(Z)$: boosted one string

Perform: string creation along z axis \mapsto perpendicular rotation \mapsto boost \mapsto backward rotation

i.e.
$$h(Z) = \mathcal{R}_{\{0, -\frac{\pi}{2}, 0\}} \mathcal{B}_{\{w\}} \mathcal{R}_{\{0, \frac{\pi}{2}, 0\}} \mathcal{S}_{\{\delta\}} Z$$

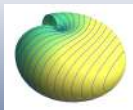
Initial data for geodesics:



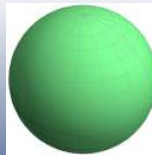
$$\mathcal{M}^+, t_i^+ = \text{const}$$



$$\mathcal{M}^-$$



$$\mathcal{M}^+$$

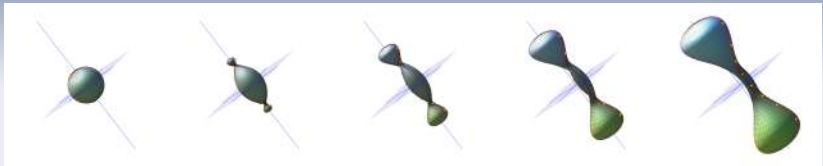


$$\mathcal{M}^-, t_i^- = \text{const}$$



Boosted one string: geodesic motion

Initial data: static in front of the impulse and emerging at $t_i^- = \text{const}$



$$t = t_i^-$$

$$t = t_i^- + 0.4$$

$$t = t_i^- + 0.8$$

$$t = t_i^- + 1.2$$

$$t = t_i^- + 1.6$$



$$t = t_i^-$$

$$t = t_i^- + 0.4$$

$$t = t_i^- + 0.8$$

$$t = t_i^- + 1.2$$

$$t = t_i^- + 1.6$$

- the test observers gain nontrivial velocities
- asymmetry induced by the string motion in the x direction

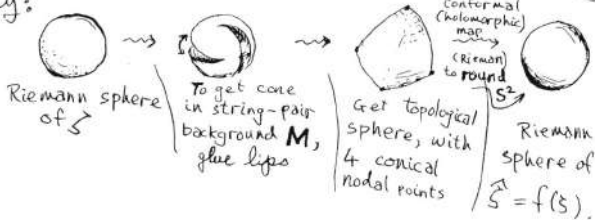


Specific function $h(Z)$: 2 strings

Y. Nutku, R. Penrose: *On impulsive gravitational waves*, Twistor Newsletter **34** (1992) 9–12

A more involved situation is provided by a pair of strings that collide

Here the function h is obtained in the following way:



It seems to be hard to find f explicitly for this case, but a Riemann theorem ensures that f exists.



Specific function $h(Z)$: 2 strings

J. Podolský, J.B. Griffiths: *The collision and snapping of cosmic strings generating spherical impulsive gravitational waves*, *Class. Quant. Grav.* **17** (2000) 1401–13

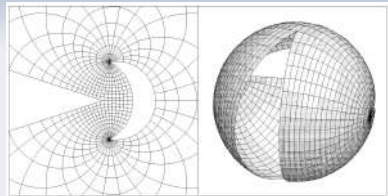
Two cosmic strings:

Located: along y and z

$$h(Z) = \left(\frac{iZ^{1-\delta} - 1}{Z^{1-\delta} - i} \right)^{1-\varepsilon}$$

Problem: $h(Z) \neq Z$ for $\delta = 0 = \varepsilon$

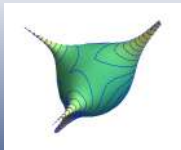
Solution: additional rotation



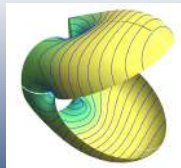
Initial data:



$\mathcal{M}^+, t_i^+ = \text{const}$



\mathcal{M}^-



\mathcal{M}^+

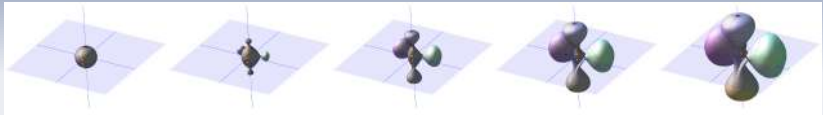


$\mathcal{M}^-, t_i^- = \text{const}$



Two strings: geodesic motion

Initial data: static *in front of* the impulse and emerging at $t_i^- = \text{const}$



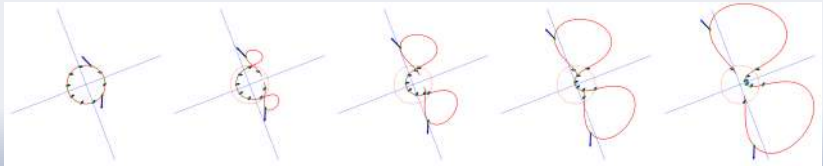
$$t = t_i^-$$

$$t = t_i^- + 0.4$$

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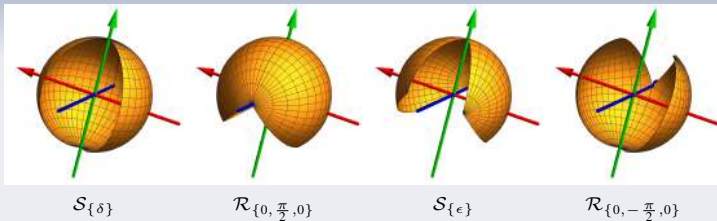
- test observers dragged by the ends of strings
- asymmetry: different values of deficit angles and ‘non-penpendicular’ strings positions



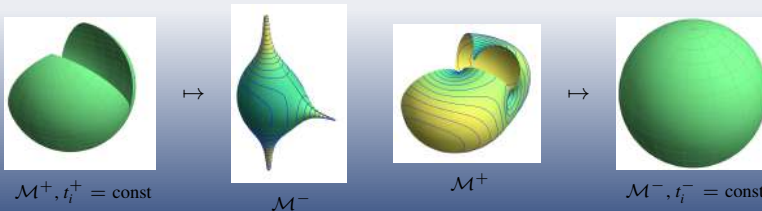
Two strings: peculiar case

Perform: string \mapsto perpendicular rotation \mapsto string \mapsto backward rotation

i.e.
$$h(Z) = \mathcal{R}_{\{0, -\frac{\pi}{2}, 0\}} \mathcal{S}_{\{\epsilon\}} \mathcal{R}_{\{0, \frac{\pi}{2}, 0\}} \mathcal{S}_{\{\delta\}} Z$$



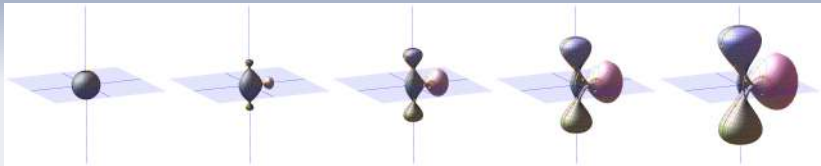
Initial data:





Peculiar two string case: geodesic motion

Initial data: static in front of the impulse and emerging at $t_i^- = \text{const}$



$t = 0$

$t = t_i^- + 0.4$

$t = t_i^- + 0.8$

$t = t_i^- + 1.2$

$t = t_i^- + 1.6$



$t = 0$

$t = t_i^- + 0.4$

$t = t_i^- + 0.8$

$t = t_i^- + 1.2$

$t = t_i^- + 1.6$

- test observers dragged by the ends of strings
- asymmetry: different values of deficit angles and only *three* string pieces



Our contribution

We were mainly interested in:

- *expanding* impulsive gravitational waves propagating on a flat background
- construction of specific complex mappings $h(Z)$ encoding the source properties
- geometric description of resulting spacetimes
- effects of the expanding impulse on free test particles

For more detail see:

- J. Podolský, R. Švarc: *Refraction of geodesics by impulsive spherical gravitational waves in constant-curvature spacetimes with a cosmological constant*, Phys. Rev. D **81** (2010) 124035
- M. Karamazov: *Impulsive gravitational waves*, Bachelor Thesis, Charles University (2015)
- J. Podolský, C. Sämann, R. Steinbauer, R. Švarc: *The global uniqueness and C^1 -regularity of geodesics in expanding impulsive gravitational waves*, Class. Quantum Grav. **33** (2016) 195010
- D. Kofroň, M. Karamazov, R. Švarc: *An interpretation of spacetimes with expanding gravitational impulses generated by a pair of snapped cosmic strings* to be submitted



... thank you for your attention ...