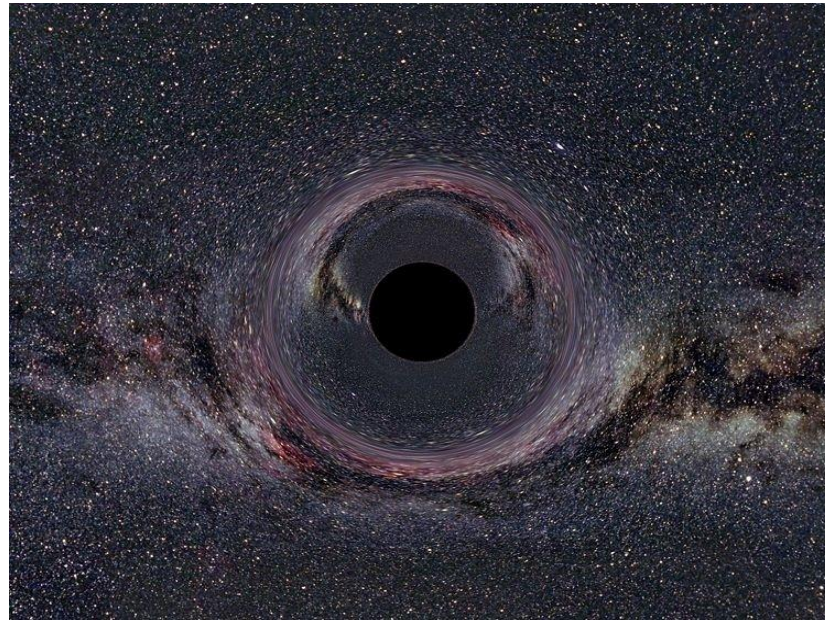


# Killing tensors: from spinning tops to SUSY in the sky

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**Selected topics in General Relativity 2024/25**

# Plan for the course

## Lecture 1: *Introduction to hidden symmetries*

- a) Why hidden? Basics of symplectic geometry
- b) Symmetries in GR: Killing & KY tensors

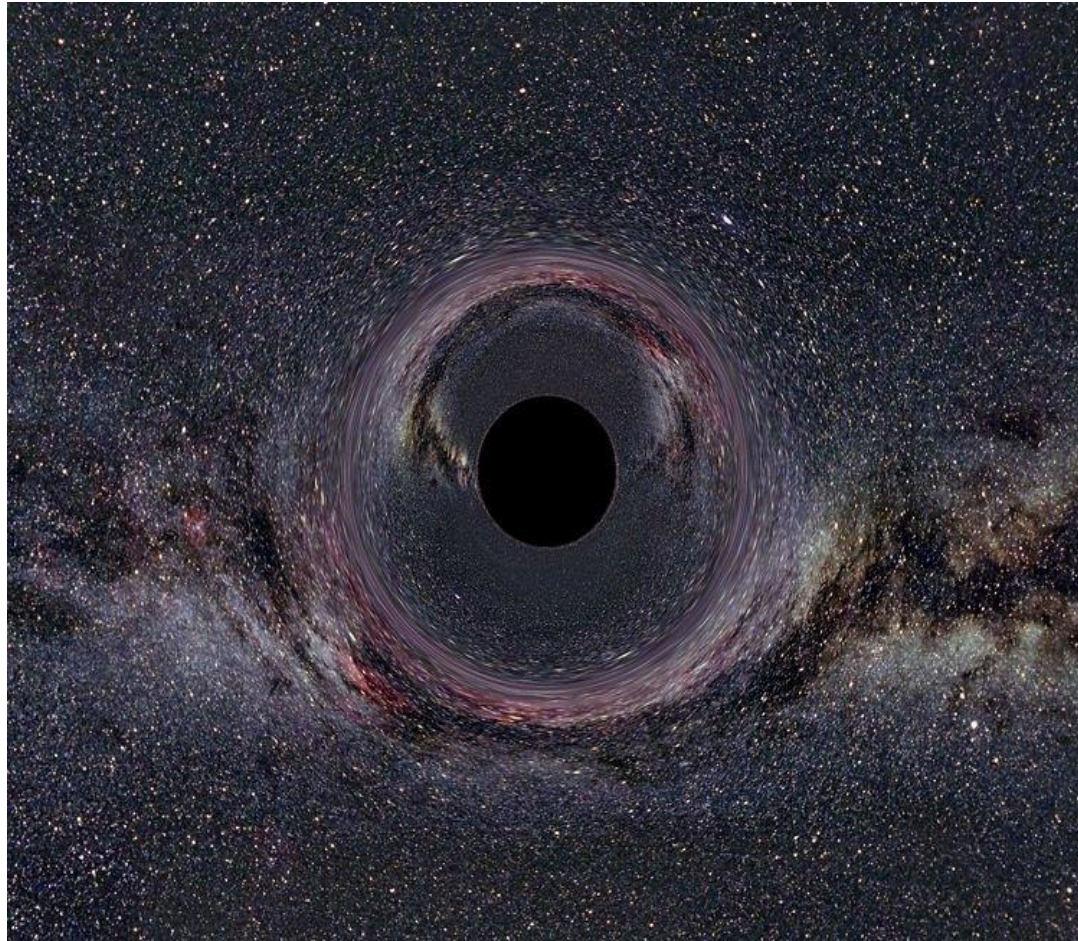
## Lecture 2: *Lifting mechanical systems & their symmetries*

- a) More on Killing and KY tensors
- b) Eisenhart lift & other geometrizations of mechanics
- c) From spinning tops to spacetimes with higher-rank KT

## Lecture 3: *SUSY in the sky*

- a) More on Killing-Yano tensors
- b) Classical spinning particle & its “fake SUSY”
- c) Connections with other descriptions of spin

# Lecture 1: Introduction to hidden symmetries

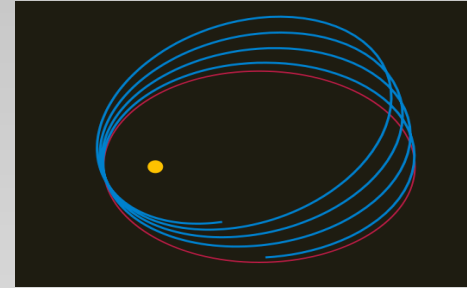


I) Why hidden? Basics of  
symplectic geometry

# Laplace-Runge-Lenz vector

Central force:

$$E, \vec{L}$$

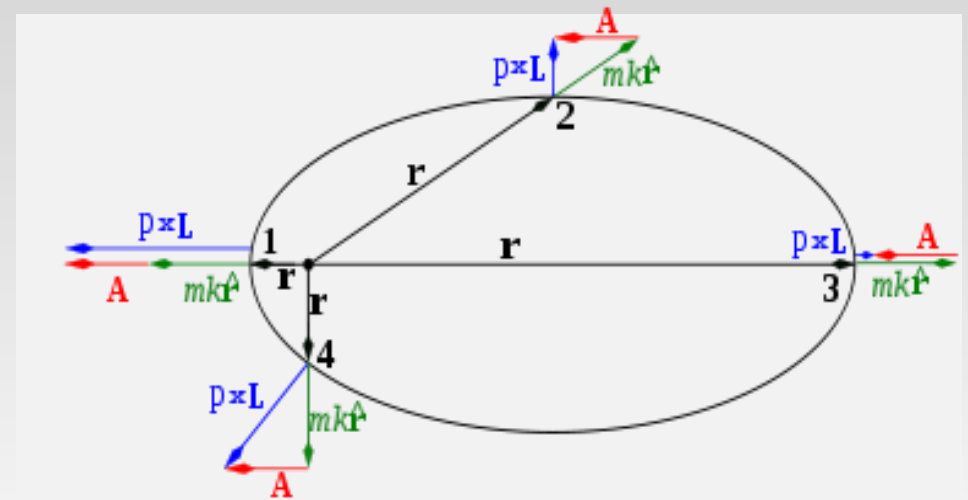


Kepler problem:

$$\vec{F} = -\frac{k}{r^2} \hat{r}$$

$$\vec{A} = \vec{p} \times \vec{L} - mk\hat{r}$$

Laplace-Runge-Lenz vector



Wikipedia

- motion maximally superintegrable

$$\vec{A} \cdot \vec{L} = 0$$

$$A^2 = m^2 k^2 + 2mEL^2$$

...hidden (dynamical) symmetry

# Symplectic 2-form

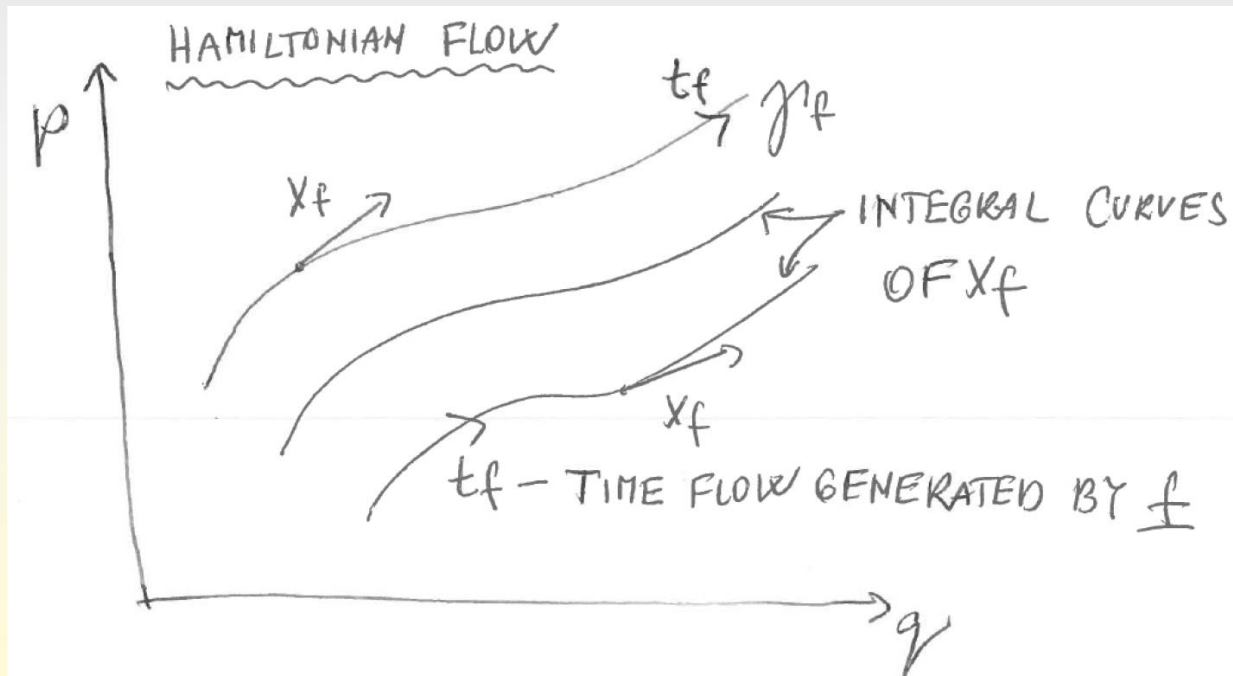
- Symplectic manifold:  $\omega = \frac{1}{2} \omega_{AB} d\xi^A \wedge d\xi^B$  (non-degenerate, closed 2-form)

- **Hamiltonian vector field** generated by function  $f$ :

$$X_f^A = \Omega^{AB} \partial_B f$$

- **Poisson bracket**

$$\{f, g\} = df \cdot \Omega \cdot dg = -\omega(X_f, X_g)$$



$$X_f = \{\cdot, f\} = \frac{d}{dt_f}$$

# Darboux theorem

- There exist Darboux coordinates:  $\xi^A = (q^i, p_j)$  s.t

$$\Omega^{\alpha\beta} = \begin{pmatrix} q^i & p_i \\ p_i & \begin{array}{c|c} \begin{matrix} 0 & 1 \\ -1 & 0 \end{matrix} & \emptyset \\ \hline \emptyset & \begin{matrix} 0 & 1 \\ -1 & 0 \end{matrix} \\ \hline \emptyset & \emptyset & \begin{matrix} 0 & 1 \\ -1 & 0 \end{matrix} \end{array} \end{pmatrix} \Leftrightarrow \omega_{\alpha\beta} = \begin{pmatrix} \begin{array}{c|c} \begin{matrix} 0 & -1 \\ 1 & 0 \end{matrix} & \emptyset \\ \hline \emptyset & \begin{matrix} 0 & -1 \\ 1 & 0 \end{matrix} \end{array} \end{pmatrix}$$

$$\omega = dp_i \wedge dq^i$$

- Poisson brackets then take the standard form:

$$\{f, g\} = \frac{\partial f}{\partial q^i} \frac{\partial g}{\partial p_i} - \frac{\partial f}{\partial p_i} \frac{\partial g}{\partial q^i}$$

# Canonical equations

- Time evolution: generated by the **Hamiltonian H**

$$X_H = \{\cdot, H\} = \frac{\partial H}{\partial p_j} \frac{\partial}{\partial q^j} - \frac{\partial H}{\partial q^j} \frac{\partial}{\partial p_j}$$

its **integral curves**  $\gamma(t) = (p_j(t), q^j(t))$

$$X_H = \frac{d}{dt} = \frac{dq^j}{dt} \frac{\partial}{\partial q^j} + \frac{dp_j}{dt} \frac{\partial}{\partial p_j}$$

which yields Hamilton's canonical equations

$$\dot{q}^j = \frac{\partial H}{\partial p_j}, \quad \dot{p}_j = -\frac{\partial H}{\partial q^j}.$$

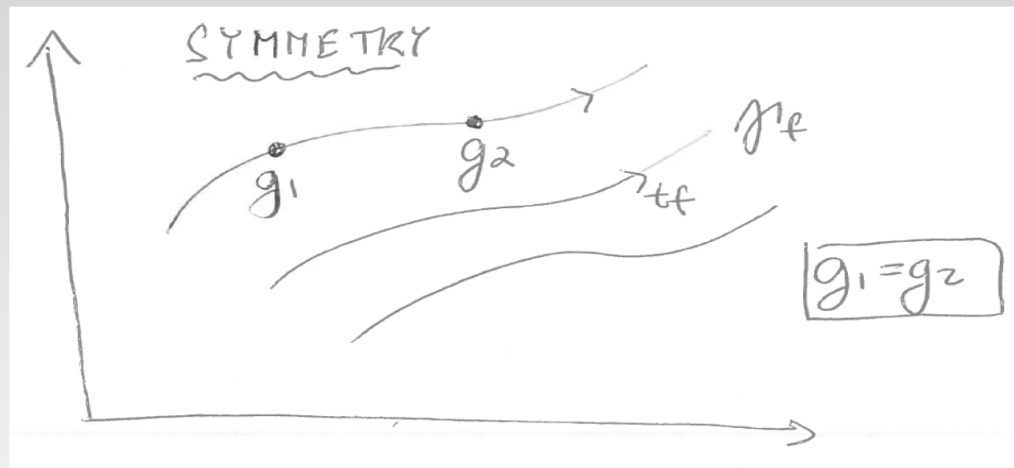
- more generally:  $\dot{f} = \frac{df}{dt} = X_H \cdot df = \{f, H\}$



# Symmetry in phase space

- Symmetry of a phase space function  $g$ :

$$\mathcal{L}_{X_f} g = X_f(g) = X_f \cdot dg = \frac{dg}{dt_f} = \{g, f\} = 0$$



- Spec:  $f=H$ , then  $g$  is preserved during the time evolution

$$\{g, H\} = 0$$

...  $g$  is an **integral of motion**

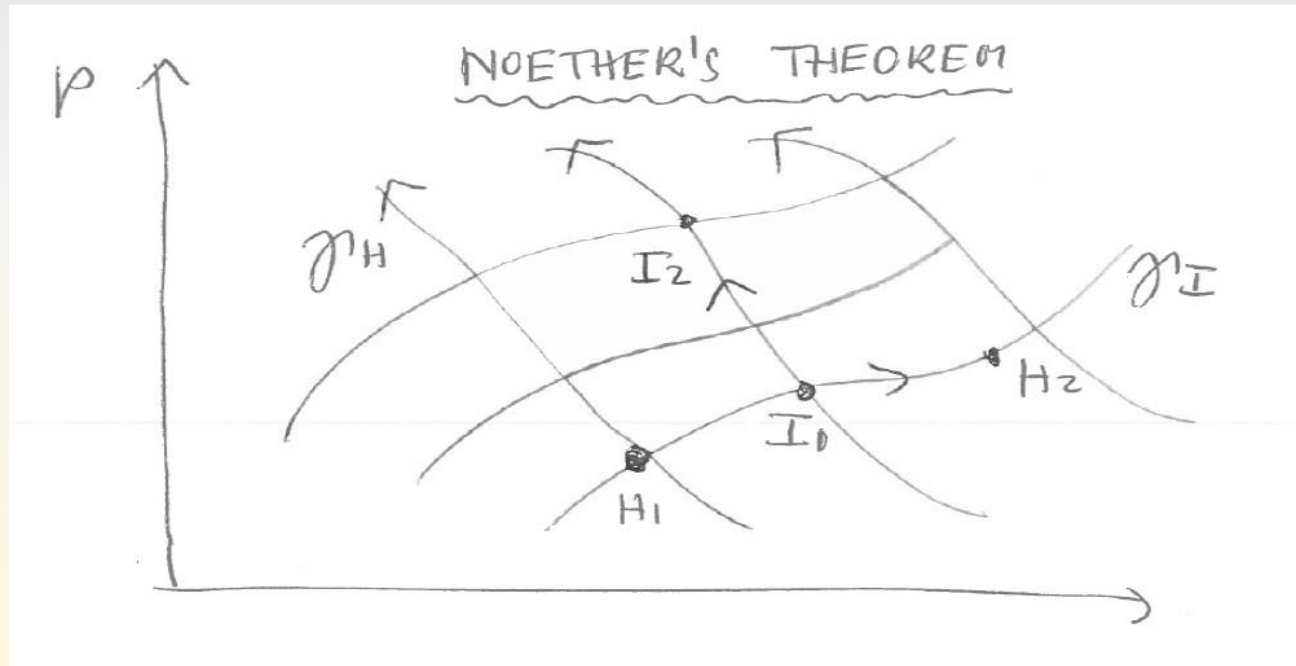
# Noether's theorem

**Noether theorem (Version 4: Hamiltonian version).** Let there be a symmetry of the Hamiltonian generated by the Hamiltonian vector field  $X_I$ . Then  $I$  is an integral of motion.

proof:

$$\{H, I\} = \frac{dH}{dt_I} = 0 \quad \Rightarrow \quad \{I, H\} = \frac{dI}{dt} = 0$$

- Geometrical meaning is non-trivial:



# Gauge freedom of Hamiltonian dynamics

- Encoded in **Canonical transformations**:

$$Q^j = Q^j(q^i, p_i), \quad P_j = P_j(q^i, p_i)$$

such that

$$\omega = dp_i \wedge dq^i = dP_i \wedge dQ^i$$

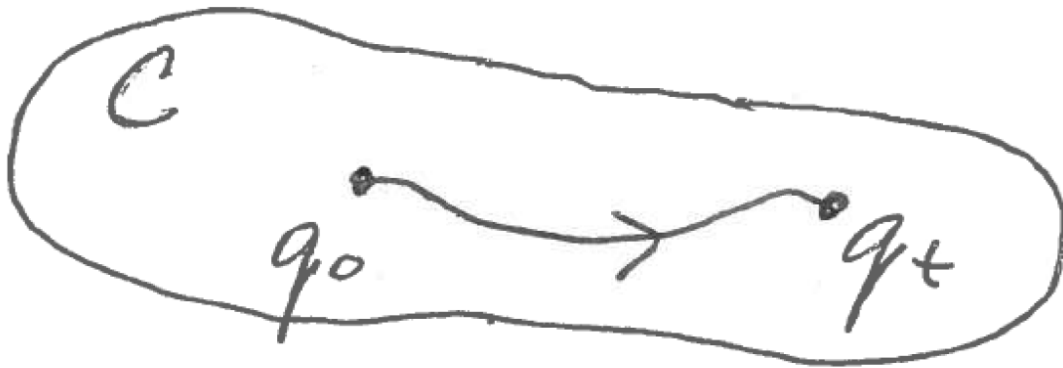
(form of Hamilton's equations remain the same)

**One cannot distinguish p's from q's!!!**

# Gauge freedom of Hamiltonian dynamics

- Spec: *Phase space* = cotangent bundle  $T^*C$

Configuration space  $C$ :



- $C$  is a manifold  $M$

$$(q^1, \dots, q^n)$$

- Typically equipped with the **metric**

- There exists a **canonical projection**:

$$\pi : T^*(\mathcal{M}) \rightarrow \mathcal{M}$$

$$\pi : (q, p) \rightarrow q$$

**breaks the symmetry!**

- Induces a map on vectors

$$\tilde{X}_I = \pi^*(X_I) = \pi^* \left( \frac{\partial I}{\partial p_i} \frac{\partial}{\partial q_i} - \underbrace{\frac{\partial I}{\partial q^i} \frac{\partial}{\partial p_i}}_{\text{kills these directions}} \right) = \frac{\partial I}{\partial p_i} \frac{\partial}{\partial q_i}$$

kills these directions

# Hidden (dynamical) symmetries

Noether theorem (Version 4: Hamiltonian version). Let there be a symmetry of the Hamiltonian generated by the Hamiltonian vector field  $X_I$ . Then  $I$  is an integral of motion.

Can distinguish isometries from dynamical symmetries:

$$\pi^*(X_I) = \begin{cases} \text{vector field on } C & \underline{\text{isometry}} \\ \text{not well defined on } C & \underline{\text{dynamical (hidden) symmetry}} \end{cases}$$

The latter is a genuine symmetry of the phase space and does not have well defined geometric meaning on  $C$ .

- **Trivial example of hidden symmetry:** energy

$$E = H = \frac{p^2}{2m} \quad \Rightarrow \quad X_E = \{\cdot, E\} = \frac{p}{m} \frac{\partial}{\partial q} = \pi^*(X_E)$$

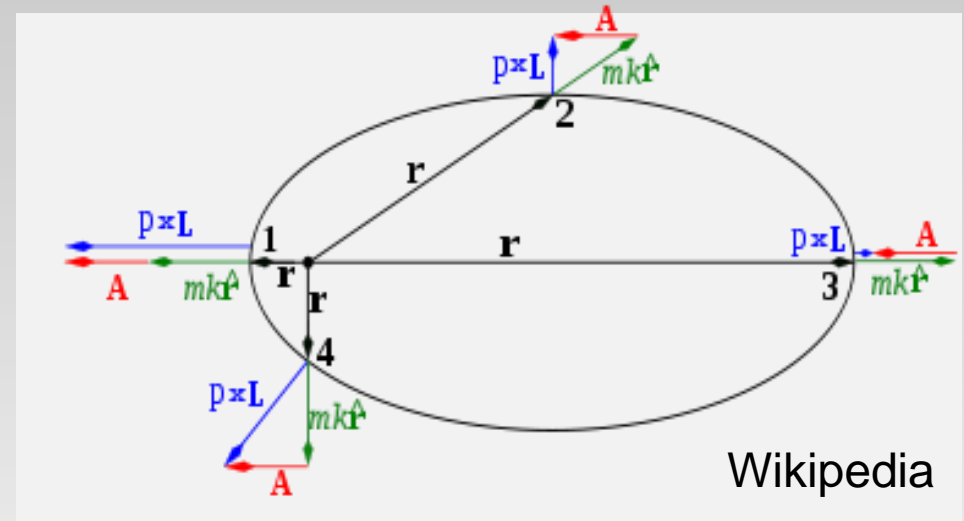
... no transformation of  $C$  in a given instant of time indicates that  $E$  should be conserved (time invariance of  $H$ )

# Hidden (dynamical) symmetries

- Non-trivial example: Laplace-Runge-Lenz vector

$$\vec{F} = -\frac{k}{r^2} \hat{r}$$

$$\vec{A} = \vec{p} \times \vec{L} - mk\hat{r}$$



- Corresponding vector flow: **(Exercise)**

$$X_{A^i} = (2x^i p^k - \delta_k^i x \cdot p - p^i x^k) \frac{\partial}{\partial x^k} - \left( \delta_k^i p^2 - p^i p_k - mk \delta_k^i \frac{1}{r} + mk \frac{x^i x^k}{r^3} \right) \frac{\partial}{\partial p^k}$$

- Canonical projection yields:

$$\pi^*(X_{A^i}) = (2x^i p^k - \delta_k^i x \cdot p - p^i x^k) \frac{\partial}{\partial x^k}$$

**Hidden  
symmetry**

## II) Symmetries in GR

# Symmmeries in GR

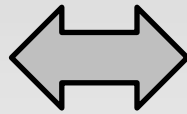
Particle motion  
(geodesics)

$$H = \frac{1}{2} g^{\mu\nu} p_\mu p_\nu$$

$$p^\mu \nabla_\mu p^\nu = 0$$

Linear in momentum conserved quantities:

$$C_K = K^\mu p_\mu$$



$$\nabla_{(\mu} K_{\nu)} = 0$$

...Killing vector equation

Proof:

$$\dot{C}_K = p^\nu \nabla_\nu (K^\mu p_\mu) = p^\nu p^\mu \nabla_{(\nu} K_{\mu)} + K^\mu \underbrace{p^\nu \nabla_\nu p_\mu}_0 = 0$$

Hamiltonian vector field:

$$X_{C_K} = K^\mu \frac{\partial}{\partial x^\mu} - \frac{\partial K^\lambda}{\partial x^\mu} p_\lambda \frac{\partial}{\partial p_\mu}$$

$$\pi_* (X_{C_K}) = K^\mu \frac{\partial}{\partial x^\mu} = K$$

**...isometry**

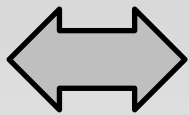
$$\mathcal{L}_\xi g = 0$$



## Higher-order conserved quantities

$$C_K = K^{\mu_1 \dots \mu_p} p_{\mu_1} \dots p_{\mu_p}$$

(Exercise)



$$\nabla_{(\mu} K_{\nu_1 \dots \nu_p)} = 0$$

...Killing tensor equation

Walker & Penrose, Comm. Math. Phys. 18 ,  
265 (1970). (Stackel 1895).

Hamiltonian vector field:

$$\pi_* (X_{C_K}) = p K^{\mu_1 \dots \mu_{p-1} \nu} p_{\mu_1} \dots p_{\mu_{p-1}} \frac{\partial}{\partial x^\nu}$$

**...dynamical symmetry**

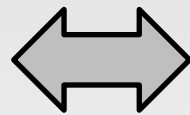
# Parallel transport

$$w^\mu = f^{\mu\nu} p_\nu$$

Impose that:

$$p^\mu \nabla_\mu w^\nu = 0$$

$$w \cdot p = 0$$



**(Exercise)**

$$\nabla_{(\kappa} f_{\mu)\nu} = 0$$

$$f_{\mu\nu} = f_{[\mu\nu]}$$

**...Killing-Yano equation**

K. Yano, Ann. Math. 55, 328 (1952). [Kashiwada 1968 & Tachibana 1969]  
R. Penrose, Ann. N.Y. Acad. Sci. , 125 (1973).

# Summary of lecture 1

- 1) **Hidden (dynamical) symmetries** are genuine symmetries of the phase space – they are “invisible” on the (preferred) configuration space.
- 2) In GR, hidden symmetries described by:
  - (symmetric) **Killing** tensors, and
  - (antisymmetric) **Killing-Yano** tensors.

Although we derived these as symmetries of the particle motion, they have far-reaching consequences for the properties of the spacetime and the dynamics of fields in it.

- 3) The latter are **more fundamental** – they are a “square root” of Killing tensors

$$K_{\mu\nu} = f_{\mu\alpha} f_{\nu}^{\alpha}$$

**(Exercise)**