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Symmetries of de Sitter and anti-de Sitter space-times are discussed. In particular, Lorentz boosts, which appear to be given previously only in the infinitesimal form, are presented here in closed finite forms.

1. Introduction

It is well-known that the maximal number of isometries of a four-dimensional space-time is 10 corresponding to the same number of independent Killing vectors (infinitesimal generators of the symmetry transformations). Such maximally symmetric space-times have a constant curvature R (see, e.g.[1]). According to the sign of R , there are three possibilities: one gets flat Minkowski space-time ($R = 0$), de Sitter space-time ($R > 0$) and anti-de Sitter space-time ($R < 0$). These space-times are the only conformally flat vacuum solutions of Einstein's field equations with cosmological constant λ . The constant curvature R is related to λ by $R = 4\lambda$.

2. Symmetries of de Sitter space-time

De Sitter manifold can be visualized as a hyperboloid

$$-z_0^2 + z_1^2 + z_2^2 + z_3^2 + z_4^2 = \alpha^2, \quad (1)$$

$\alpha = \sqrt{\frac{\lambda}{3}}$, embedded in a flat five-dimensional space-time

$$ds^2 = -dz_0^2 + dz_1^2 + dz_2^2 + dz_3^2 + dz_4^2. \quad (2)$$

We use coordinates (η, x, y, z) covering the entire hyperboloid given by

$$\begin{aligned}
z_0 &= \frac{1}{2\eta}(\alpha^2 + s) \\
z_1 &= \alpha \frac{x}{\eta} \\
z_2 &= \alpha \frac{y}{\eta} \\
z_3 &= \alpha \frac{z}{\eta} \\
z_4 &= \frac{1}{2\eta}(\alpha^2 - s)
\end{aligned} \tag{3}$$

where $s = -\eta^2 + x^2 + y^2 + z^2$ and $\eta, x, y, z \in (-\infty, +\infty)$. In the global coordinates (3) de Sitter metric can be written as

$$ds^2 = \frac{\alpha^2}{\eta^2}(-d\eta^2 + dx^2 + dy^2 + dz^2). \tag{4}$$

This conformally flat metric (see, e.g. [3]) is related to the ‘standard’ form of de Sitter metric

$$ds^2 = -dt^2 + \exp\left(2\frac{t}{\alpha}\right)(dx^2 + dy^2 + dz^2) \tag{5}$$

by a simple transformation $\eta = \alpha \exp(-t/\alpha)$. However, the metric (5) covers only one half ($z_0 + z_4 > 0$) of the manifold (1). Thus, one needs another coordinate chart of the form (5) to cover also $z_0 + z_4 < 0$; such a chart is given by $\eta = -\alpha \exp(-t^*/\alpha)$. We prefer coordinates (3) since they do not only cover the entire manifold but the symmetries can be seen most easily just in these coordinates.

Killing vectors are (see [2],[3], where, however, the expressions are given in ‘standard’ coordinates used in (5) and thus have more complicated forms):

$$\begin{aligned}
\xi_{(1)}^\mu &= (0, 1, 0, 0) \\
\xi_{(2)}^\mu &= (0, 0, 1, 0) \\
\xi_{(3)}^\mu &= (0, 0, 0, 1) \\
\xi_{(4)}^\mu &= (0, y, -x, 0) \\
\xi_{(5)}^\mu &= (0, z, 0, -x) \\
\xi_{(6)}^\mu &= (0, 0, z, -y) \\
\xi_{(7)}^\mu &= (\eta, x, y, z) \\
\xi_{(8)}^\mu &= -2x\left(\eta, x - \frac{s}{2x}, y, z\right) \\
\xi_{(9)}^\mu &= -2y\left(\eta, x, y - \frac{s}{2y}, z\right) \\
\xi_{(10)}^\mu &= -2z\left(\eta, x, y, z - \frac{s}{2z}\right).
\end{aligned} \tag{6}$$

By integrating differential equations [2] of the forms

$$\left. \frac{\partial x'^{\mu}}{\partial \varepsilon} \right|_{\varepsilon=0} = \xi^{\mu}(x^{\alpha}) \quad (7)$$

we find finite symmetry transformations. Clearly, there are 3 spatial translations and 3 spatial rotations of the form exactly the same as in the flat Minkowski space-time. There is one time translation accompanied by a spatial dilatation given by

$$\begin{aligned} \eta &\rightarrow \eta' = (A + 1)\eta \\ x &\rightarrow x' = (A + 1)x \\ y &\rightarrow y' = (A + 1)y \\ z &\rightarrow z' = (A + 1)z \end{aligned} \quad (8)$$

where $A = \text{const}$. Remaining 3 symmetries are the analogues of Lorentz boosts. The boost in x -direction is given by

$$\begin{aligned} \eta &\rightarrow \eta' = \frac{\eta}{sB^2 + 2xB + 1} \\ x &\rightarrow x' = \frac{sB + x}{sB^2 + 2xB + 1} \\ y &\rightarrow y' = \frac{y}{sB^2 + 2xB + 1} \\ z &\rightarrow z' = \frac{z}{sB^2 + 2xB + 1} \end{aligned} \quad (9)$$

where $B = \text{const}$. The boosts in y , z -directions can be obtained from (9) simply by a cyclic permutation $x \rightarrow y \rightarrow z \rightarrow x$.

3. Symmetries of anti-de Sitter space-time

The anti-de Sitter manifold can be visualized as a hyperboloid

$$-z_0^2 - z_1^2 + z_2^2 + z_3^2 + z_4^2 = -\beta^2, \quad (10)$$

$\beta = \sqrt{-\frac{\lambda}{3}}$, embedded in a flat five-dimensional space-time

$$ds^2 = -dz_0^2 - dz_1^2 + dz_2^2 + dz_3^2 + dz_4^2. \quad (11)$$

We use coordinates (η, x, y, z) covering the entire hyperboloid given by

$$\begin{aligned} z_0 &= \frac{1}{2x}(\beta^2 + s) \\ z_1 &= \beta \frac{\eta}{x} \\ z_2 &= \beta \frac{y}{x} \\ z_3 &= \beta \frac{z}{x} \\ z_4 &= \frac{1}{2x}(\beta^2 - s) \end{aligned} \quad (12)$$

where $\eta, x, y, z \in (-\infty, +\infty)$. In the global coordinates (12) anti-de Sitter metric can be written as

$$ds^2 = \frac{\beta^2}{x^2}(-d\eta^2 + dx^2 + dy^2 + dz^2). \quad (13)$$

This conformally flat metric is related to the ‘standard’ form (see, e.g. [1]) of anti-de Sitter metric

$$ds^2 = -\cosh^2 r dt^2 + \beta^2 dr^2 + \beta^2 \sinh^2 r (d\vartheta^2 + \sin^2 \vartheta d\varphi^2) \quad (14)$$

by a transformation

$$\begin{aligned} \eta &= X \cosh r \cos\left(\frac{t}{\beta}\right) \\ x &= X \\ y &= X \sinh r \cos \vartheta \\ z &= X \sinh r \sin \vartheta \cos \varphi \end{aligned} \quad (15)$$

where $X = \beta/(\cosh r \sin(\frac{t}{\beta}) + \sinh r \sin \vartheta \sin \varphi)$.

In the coordinates (12) Killing vectors have the following form

$$\begin{aligned} \xi_{(1)}^\mu &= (1, 0, 0, 0) \\ \xi_{(4)}^\mu &= (y, 0, \eta, 0) \\ \xi_{(5)}^\mu &= (z, 0, 0, \eta) \\ \xi_{(8)}^\mu &= 2\eta\left(\eta + \frac{s}{2\eta}, x, y, z\right), \end{aligned} \quad (16)$$

six remaining vectors having the same form as in (6). Clearly, there is one time translation (in η direction), two spatial translations (in y, z -directions), two ‘Minkowskian’ boosts (in $\eta - y$ and $\eta - z$ planes) and one spatial rotation (in $y - z$ plane). There is also the symmetry given by (8). Remaining 3 symmetries are given by

$$\begin{aligned} \eta \rightarrow \eta' &= \frac{sC + \eta}{-sC^2 - 2\eta C + 1} \\ x \rightarrow x' &= \frac{x}{-sC^2 - 2\eta C + 1} \\ y \rightarrow y' &= \frac{y}{-sC^2 - 2\eta C + 1} \\ z \rightarrow z' &= \frac{z}{-sC^2 - 2\eta C + 1} \end{aligned} \quad (17)$$

where $C = \text{const}$,

$$\begin{aligned}
 \eta &\rightarrow \eta' = \frac{\eta}{sD^2 + 2yD + 1} \\
 x &\rightarrow x' = \frac{x}{sD^2 + 2yD + 1} \\
 y &\rightarrow y' = \frac{sD + y}{sD^2 + 2yD + 1} \\
 z &\rightarrow z' = \frac{z}{sD^2 + 2yD + 1}
 \end{aligned}
 \tag{18}$$

where $D = \text{const}$ and analogously (18) with $y \leftrightarrow z$. This completes the list of all isometries of de Sitter and anti-de Sitter space-times.

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References

- [1] Hawking S.W. and Ellis G.F.R.: *The Large Scale Structure of Space-time*, Cambridge University Press, Cambridge, 1973.
- [2] Robertson H.P. and Noonan T.W.: *Relativity and Cosmology*, W.B. Saunders, Philadelphia, 1969. ■
- [3] Schmidt H.-J.: *Potsdam preprint PRE-ZIAP 91-04*, 1991; submitted to Fortschr. Phys.